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Nishitsuji, Yohei; Ruigrok, E; Gomez, M; Wapenaar, Kees; Draganov, Deyan

DOI 10.1190/INT-2015-0225.1

Publication date 2016 **Document Version** Accepted author manuscript

Published in Interpretation

Citation (APA) Nishitsuji, Y., Ruigrok, E., Gomez, M., Wapenaar, K., & Draganov, D. (2016). Reflection imaging of aseismic zones of the Nazca slab by global-phase seismic interferometry. *Interpretation*, *4*(3), SJ1-SJ16. https://doi.org/10.1190/INT-2015-0225.1

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Full title: Reflection Imaging of Aseismic Zones of the Nazca slab by Global-phase Seismic Interferometry

Author's names: Yohei Nishitsuji^{1*}, Elmer Ruigrok², Martín Gomez³, Kees Wapenaar¹, Deyan Draganov¹

Affiliation: ¹Department of Geoscience and Engineering, Delft University of Technology, Delft, The Netherlands
²Department of Earth Sciences, Utrecht University, Utrecht, The Netherlands; R&D Seismology and Acoustics, Royal Netherlands Meteorological Institute (KNMI), De Bit, The Netherlands
³International Center for Earth Sciences, Comision Nacional de Energia Atomica, Buenos Aires, Argentina

Date of submission: 30-Dec, 2015

Abbreviated title: Reflection imaging of aseismic zones

Corresponding author:

name : Yohei Nishitsuji

address: Department of Geoscience and Engineering,

Delft University of Technology

Stevinweg 1, 2628 CN Delft, Netherlands

P.O. Box 5048, 2600 GA Delft, Netherlands

email : <u>y.nishitsuji@tudelft.nl</u>

1 Abstract

2	Obtaining detailed images of aseismic parts of subducting slabs remains a large
3	challenge for understanding slab dynamics. Hypocenter mapping cannot be used for
4	the purpose due to the absence of seismicity, while the use of receiver functions might
5	be compromised by the presence of melt. Global tomography can be used to identify
6	the presence of the slab, but does not reveal its structure in detail. Here, we show how
7	detailed images can be obtained using global-phase seismic interferometry. The
8	method provides high-resolution (< 15 km in depth) pseudo zero-offset (i.e., co-located
9	source and receiver) reflection information. We apply the method to aseismic zones of
10	the Nazca slab where initiation of possible slab tearing and plume decapitation are
11	identified by global tomography and electrical conductivity, respectively. We obtain an
12	image of the Moho and the mantle, and find an attenuated area in the image consistent
13	with the presence of an aseismic dipping subducting slab. However, the interpretation is
14	not unambiguous. The results confirm that the method is useful for imaging aseismic
15	transects of slabs.

17 **INTRODUCTION**

18	It has been shown that at the northern part of Central Chile (30 - 33°S) the
19	Nazca slab is of the flat type (e.g., Rosenbaum et al., 2005; Anderson et al., 2007;
20	Eakin et al., 2014). At that part, the upwelling plume was recently imaged (Booker et
21	al., 2004). Still, the slab's geometry in the southern part of central Chile (34 - 37°S) is
22	unclear and it is unknown whether that part of the slab is not torn (e.g., Gilbert et al.,
23	2006; Pesicek et al., 2012).
24	One of the challenges in imaging the slab in this region by seismological
25	methods relates to the absence of seismicity. Although hypocenter mapping is a useful
26	method for identifying the Wadati-Benioff zone (e.g., Cahill and Isacks, 1992;
27	Syracuse and Abers, 2009; Bloch et al., 2014), it cannot be used to image the aseismic
28	region.
29	The receiver-function method (e.g., Langston 1979; Audet et al., 2009;
30	Kawakatsu and Yoshioka, 2011) can be used to image aseismic regions, but so far has
31	not yielded images of the aseismic zone in this region. Yuan et al. (2000) suggest that
32	the reason for this might be the possible completion of the gabbro-eclogite

33	transformation within the Nazca slab. Gilbert et al. (2006) suggest large attenuation of
34	S-wave energy in the mantle wedge as another possible reason.
35	Global tomography (e.g., Aki et al., 1977; Dziewonski et al., 1977; Boschi
36	and Becker, 2011) is a tool for investigating global-scale geodynamics and it can be
37	used for imaging aseismic zones. However, the method's resolution (\approx 50 km) poses
38	limitations on estimating the slab's exact location and continuity at local scale, thus
39	leaves a lot of uncertainties.
40	The reflection method with active sources (explosives, vibroseis, airguns)
41	provides the needed high-resolution imaging capabilities, but its depth penetration is
42	fundamentally limited by the strength of the used sources.
43	Here, we demonstrate the usefulness of an alternative seismic technique to
44	image the aseismic slab zone with high resolution, namely seismic interferometry (SI)
45	for body-wave retrieval (e.g., Claerbout, 1968; Scherbaum, 1987a,b; Daneshvar et al.,
46	1995; Wapenaar, 2003) using global phases (GloPSI) (Ruigrok and Wapenaar, 2012).
47	Global phases are seismic phases that travel through the Earth's core before reaching
48	the surface. They are induced by earthquakes at epicentral distances greater than 120°

49	(global distances). The global phases are extracted from the continuous field
50	recordings and used as contributions from separate transient sources. For the
51	considered configuration, this is closely related to work of Kumar and Bostock (2006)
52	and Nowack et al. (2007). For a horizontally layered (1D) acoustic medium, SI
53	retrieves the reflection response of the medium from the autocorrelation of the
54	medium's plane-wave transmission response measured at the surface (Claerbout,
55	1968). GloPSI is a 3D generalization of the mentioned 1D case - it extends the
56	illumination to include a range of ray parameters (horizontal slownesses) allowing
57	retrieval of reflections from 3D structures. At seismic stations, these extra ray
58	parameters would come from recorded global P-wave arrivals, such as the phases PKP,
59	PKiKP, and PKIKP. These arrivals (phases) have ray parameters lower than 0.04 s/km
60	and are characterized in the mantle by nearly planar wavefronts. This makes these
61	phases suitable for SI by autocorrelation. Due to the autocorrelation, GloPSI retrieves
62	pseudo zero-offset reflection arrivals that penetrate deep enough to allow slab imaging
63	with resolution dictated by the frequency bandwidth of the phases, sensor
64	configuration and two-way traveltime difference between consecutive arrivals. GloPSI

00	may further shed light on one of the open questions in the geoscience community of
66	whether small deformations and/or detachments (< 25 km) in the slab are actually
67	present (Wortel and Spakman, 2000).
68	In the following, we show how to apply GloPSI to field waveform data. First
69	we describe the GloPSI method, then we describe the data we use, phase extraction and
70	preparation, and then we show our results and their interpretation. Our results image
71	the aseismic zone of the slab and possible deformation in the slab.
72	
72 73	<u>Global-phase seismic interferometry (GloPSI)</u>
72 73 74	<u>Global-phase seismic interferometry (GloPSI)</u> <u>Theory</u>
72 73 74 75	Global-phase seismic interferometry (GloPSI) Theory The 1D theory from Claerbout (1968) was generalized for a 3D
72 73 74 75 76	Global-phase seismic interferometry (GloPSI) Theory The 1D theory from Claerbout (1968) was generalized for a 3D inhomogeneous medium by Wapenaar (2003). Ruigrok and Wapenaar (2012) applied
72 73 74 75 76 77	Global-phase seismic interferometry (GloPSI) Theory The 1D theory from Claerbout (1968) was generalized for a 3D inhomogeneous medium by Wapenaar (2003). Ruigrok and Wapenaar (2012) applied the generalization of seismic interferometry for retrieval of body waves from the

- 79 They termed this specific application GloPSI.
- 80 The GloPSI relation for the retrieval of the zero-offset reflection response

- 81 $R(\mathbf{x}_R, \mathbf{x}_R, t)$ for co-located source and receiver at the location of station \mathbf{x}_R is
- 82 (Ruigrok and Wapenaar, 2012)

83

84
$$\sum_{\substack{P \text{ min} \\ P \text{ min} \\ \theta \text{ min}}} \sum_{\substack{\theta \text{ min} \\ \theta \text{ min}}} \left\{ T(\mathbf{x}_{R}, \mathbf{p}_{S}, -t) * T(\mathbf{x}_{R}, \mathbf{p}_{S}, t) * E_{i}(-t) * E_{i}(t) \right\} \propto (1)$$

$$\left\{ \delta(t) - R(\mathbf{x}_{R}, \mathbf{x}_{R}, -t) - R(\mathbf{x}_{R}, \mathbf{x}_{R}, t) \right\} * \overline{E}_{n}(t)$$

85

where $T(\mathbf{x}_{R},\mathbf{p}_{S},t)$ is the transmission response (selected global phase) at the receiver 86 location \mathbf{x}_{R} due to an earthquake i_{r} arriving from direction $\mathbf{p}_{s} = (p, \theta)$ with ray 87 88 parameter p and back azimuth θ , $E_i(t)$ is the source time function of the *i*-th earthquake, $\overline{E}_n(t)$ is the average of the autocorrelations of the different source time 89 90 functions, and * denotes convolution. In our case, the absolute value of the ray 91 parameter varies between 0 and 0.04 s/km, while θ varies between 0° and 360°. In 92equation (1), the summation is effectively over plane-wave sources, instead of over 93point sources. A derivation of the SI relation from point sources to plane-wave sources 94can be found in Ruigrok et al. (2010). The zero-offset reflection response retrieved by 95GloPSI can be used to image the subsurface structures in a way similar to the 96 conventional reflection seismic method with active sources. Note that GloPSI directly

97	produces zero-offset reflection responses of the subsurface, which is one of the
98	conventional goals of the active-source reflection method. With the latter, offset
99	measurements are stacked to obtain pseudo zero-offset traces (Yilmaz, 1987), as direct
100	zero-offset measurements are still commercially impractical. A difference between the
101	zero-offset section retrieved by GloPSI and an active-source pseudo zero-offset section
102	is that the virtual source in the former radiates energy vertically and near-vertically
103	down into the Earth, while in the latter the pseudo zero-offset source radiates in all
104	directions. Because of this, GloPSI will image horizontal to mildly inclined structures
105	directly, while steeply dipping structures will be manifest by a lack of reflections
106	reaching the receivers and can be interpreted by discontinuation of imaged (nearly)
107	horizontal structures. This is similar to the problem in the active-source reflection
108	method, where a steeply dipping structure lying relatively deep compared to the
109	receiver-array length, will not be imaged (e.g., Yilmaz, 1987).
110	When the length of the used receiver array is sufficiently long, relative to the
111	depth of the structure of interest, and given a sufficiently wide illumination (in terms of

112 ray parameters and back azimuths), the autocorrelation in the GloPSI relation (1) can

113	be replaced by crosscorrelation, which would permit retrieval of offset reflections as
114	well. This would allow for direct imaging of a broader range of dipping structures.
115	In Figure 1, we show in a schematic way how GloPSI would (or would not)
116	retrieve reflection responses from four different structural settings.
117	
118	Comparison with the receiver-function method
119	The receiver-function method depends on phase conversions (P-to-S or
120	S-to-P) occurring in transmission. GloPSI with P-wave phases uses reflection
121	information and depends only on the P-wave impedance contrasts, just like the
122	conventional reflection method. Comparisons of imaging results from SI and receiver
123	function have shown that SI provides images with resolution at least as high as the
124	receiver-function image (Abe et al., 2007). In cases of structural contrasts that are due
125	to relatively thin layers, SI has the potential to provide higher resolution than the
126	receiver function. For example, suppose there is a mantle structure 5 km below the
127	Moho, which is illuminated by a P -wave phase with an incidence angle of 10°. The P-
128	and S-wave velocities between the structure and the Moho are 8.1 km/s and 4.5 km/s,

129	respectively, while above the Moho the respective velocities are 5 km/s and 2.5 km/s.
130	The receivers at the surface would record the P -to- S converted waves from the two
131	boundaries with a time difference of 0.49 s – the time difference for the propagation of
132	the P- and S-waves between the mantle structure and the Moho. A virtual zero-offset
133	reflection recording, retrieved from GloPSI, would contain two P-wave reflections
134	from the impedance contrasts at the Moho and the mantle structure arriving with a time
135	difference of 1.23 s. In terms of wavelength, assuming a center frequency for both P-
136	and S-waves of 0.8 Hz, the two arrivals in the recordings used by the receiver-function
137	method would be 0.39 wavelengths apart. In the retrieved recordings from GloPSI, the
138	two P-wave reflections would be 0.99 wavelengths apart, which would allow for
139	higher resolution.
140	Thus, although until now SI or GloPSI has not been applied for imaging of
141	aseismic slab zones, these methods have the potential to image such zones with
142	temporal (depth) resolution higher than the one that can be achieved using the
143	receiver-function method.

144

145 **<u>Data</u>**

146 Study area

147Figure 2 shows the location of intermediate-depth earthquakes that have 148occurred from August 1906 to July 2014 around the Malargüe region (35.5°S), 149 Argentina. The locations are taken from the U.S. Geological Survey (USGS, 150http://earthquake.usgs.gov/earthquakes/) earthquake catalog. There could be more 151earthquakes actually present than we show in Figure 2 if they are not in the catalog. 152Note that there are no earthquakes deeper than around 200 km. There is also an 153aseismic spot beneath the Peteroa Volcano. This volcano forms part of the 154Planchón-Peteroa volcanic complex. We are interested in imaging these aseismic zones, and we achieve this using GloPSI. In Figure 2, the station GO05 of the Chilean 155156National Seismic Network and the station C02A of the Talca Seismic Network, which we use later for quality-control purpose, are also plotted. 157

158

159 MalARRgue

160 We apply GloPSI to data from the MalARRgue array (Ruigrok et al., 2012).

161	The array recorded continuously ambient noise and seismicity during 2012 in the
162	Malargüe region, Argentina, to the east of the southern part of central Chile. The array
163	consisted of a patchy subarray PV and an exploration-style 2D T-shaped subarray T
164	with arms TN and TE pointing north and east, respectively, see Figure 3. MalARRgue
165	used short-period (2-Hz) sensors borrowed from the Program for Array Seismic
166	Studies of the Continental Lithosphere (PASSCAL) managed by Incorporated
167	Research Institutions for Seismology (IRIS). The PV-array consisted of 6 irregularly
168	spaced stations labeled PV01 to PV06; the TN-array formed a line of 19 stations
169	spaced at 2 km and labeled TN02 to TN20, while the TE-array formed a line of 13
170	stations spaced at 4 km and labeled TE01 to TE13.
171	Figure 3 shows the distribution of the global earthquakes we use to extract
172	phases at the PV- and T-array, which phases are then used as input for GloPSI. The
173	T-array lies above the beginning of the Nazca's aseismic zone, where possible slab
174	tearing (Pesicek et al., 2012) and/or presence of plume decapitation (Burd et al., 2014)
175	have been proposed.

177 Selecting and extracting global phases

178	We use the vertical-component recordings of the MalARRgue array for
179	GloPSI. Using Java version of Windows Extracted from Event Data (JWEED) from
180	IRIS and a reference earthquake catalogue from USGS, from the recorded total amount
181	of global earthquakes with $M_{\rm W} \ge 5.5$, we select 66, 72, and 85 earthquakes for the PV-,
182	TN-, and TE-array, respectively (Table 1). We use PKP, PKiKP and PKIKP phases
183	(epicentral distances $\geq 120^{\circ}$), which travel through the mantle and core and arrive at
184	the stations with absolute slowness < 0.04 s/km (Kennett et al., 1995). We search the
185	phases visually using a window of 900 s, which starts 100 s before the expected arrival
186	of the specific P-wave phase; we also use as guides the phase pickings that are
187	automatically calculated by IRIS. Then, we extract the desired phases from a shorter
188	window, which is at least 200 s long. This window starts before the arrival of the
189	specific <i>P</i> -wave phase and terminates before onset of the first <i>S</i> -wave phase. Figure 4
190	shows an example of the windowing.
191	For quality control, as described below, we also use data from the station

192 GO05 from the Chilean National Seismic Network, which is situated above the seismic

zone of the Nazca slab. For GO05, we use 52 earthquakes recorded by the station
during the operation of MalARRgue (Table 1). The complete list of the used
earthquakes for MalARRgue and GO05 is given in Table 1.

196

197 Data Processing

198 Data processing for obtaining images

199 After deconvolving the recordings with the instrument response, we 200 compute power spectral densities (PSD) of the global-phase earthquakes to help us 201select a frequency band that provides adequate signal-to-noise ratio of the global 202phases. Figure 5 shows an example of the computed PSD for earthquakes of different 203magnitude higher than 5.5 that occurred at global distances. We select the band 0.3-1.0 Hz using a 5th-order butterworth filter, as in this band all signals of the earthquakes are 204 205clearly observed (Figure 5). The lower limit of our band is set at 0.3 Hz due to the 206 low-frequency limitations of the used instruments (Nishitsuji et al., 2014), as well as to 207make sure that the double-frequency microseisms noise is largely excluded.

208 After selecting the frequency band between 0.3 Hz and 1 Hz, we

209	downsample the data from the original sampling of 0.01 s to 0.25 s with the aim to
210	minimize the volume of data. After that, we normalize each selected and filtered phase
211	with respect to its maximum amplitude. We also apply despiking to trace intervals with
212	very strong (accidental) signal spikes that saturate the trace for some time (the interval
213	duration). For the TN- and TE-array, missing traces at certain stations (e.g., due to
214	despiking) are interpolated using the corresponding records at their neighboring
215	stations (Figure 6).
216	After the above preprocessing, we apply GloPSI to the selected events for
217	each of the subarrays from MalARRgue (Figure 7). The retrieved zero-offset reflection
218	trace at each station is dominated in the first few seconds by the average
219	autocorrelation convolved with a delta function, $\overline{E}_n(t) * \delta(t)$. To suppress the effect of
220	$\overline{E}_n(t)$, for each subarray we extract the effective source time functions $\overline{E}_n(t)$ from
221	each retrieved zero-offset trace per subarray for a two-way traveltime from 0 to 10 s,
222	take their mean, and subtract the mean from the individual traces in each subarray
223	(Figure 8). This does not cause any changes to signals retrieved later than 10 s, while

224 earlier than 10 s it preserves the differences between a trace and the mean. The

225 effective source time function of 10 s was selected after testing the above procedure for

values from 8 s to 13 s with steps of 1 s.

227

228 **PKP triplication**

229We also investigate the effect on our results of the PKP triplication (Adams 230and Randall, 1963) using the T-array. The PKP triplication is expected to arise for 231earthquakes at epicentral distances from about 135° to 155°. The triplicated arrivals are 232expected within 10 s from the first PKP arrival (e.g., Garcia et al., 2004). Each of the 233*PKP* triplications will contribute in the autocorrelation process to the retrieval of the 234same reflections (for example from the Moho) and thus would result in an increased 235signal-to-noise ratio of the reflections. For each transmission response, the individual 236*PKP* triplicated arrivals will also correlate with each other, which will result in the retrieval of artifacts in the result from each transmission response (cross-talk). 237238However, according to the 3D theory of SI for any inhomogeneous medium, i.e., what 239we use here, such triplication-related artifacts will cancel out after summing over the correlated transmission responses (e.g., Wapenaar, 2003). Because of this, Ruigrok and 240

241	Wapenaar (2012) suggested using global phases from a wide range of ray parameters.
242	In the summation process after the autocorrelation, this would cause the different
243	cross-talk artifacts to interact destructively. This happens, as the cross-talk artifacts
244	would be retrieved at different times. On the other hand, correlations of global phases
245	with a wide azimuthal and slowness coverage enhance the physical arrivals, i.e., the
246	signal-to-noise ratio of structures like Moho) is improved (Snieder, 2004). In our case,
247	the azimuthal coverage and the slowness variation of the earthquakes with epicentral
248	distances $\geq 120^\circ$ are sufficiently wide (see Figure 3), so we did not exclude the
249	earthquakes that would contain PKP triplications. To the contrary, if we exclude the
250	epicentral distances causing PKP triplication, only 13 earthquakes would remain for
251	both arms of the T-array from the original 72 and 85 earthquakes for the TN- and
252	TE-array, respectively. A reduced number of used earthquakes would result in
253	deterioration of the retrieved reflections from deeper structures.
254	In Figure 9, we show a comparison of the obtained images of the subsurface
255	when including and excluding the PKP triplication. When the velocity model of

256 Gilbert et al. (2006) is used for the depth conversion, the top of the Moho is interpreted

257	at a depth of 35 km, while the possible effect of the PKP triplication should be seen
258	between depths of 35 km and 66 km. The comparison of the results in Figure 9 shows
259	that the Moho in the results when earthquakes with triplications are included is well
260	imaged without apparent large-amplitude "ringing" around it due to the PKP
261	triplication. In our context, "large" means the amplitude as large as the one of the first
262	Moho reflection, i.e., the reflection at around 30 km in Figure 9. There are some slight
263	differences in the weaker-amplitude events (e.g., positive-amplitude waveforms about
264	10 km after the Moho refection), which we attribute to an insufficient integration over
265	the small number of the earthquakes (only 13) when earthquakes with triplications are
266	excluded. Note that the triplication "ringing" should be present also shallower than the
267	Moho, but there it would be suppressed, even when present, by the subtraction of the
268	averaged source time function $\overline{E}_n(t)$.
269	The same reasoning for the suppression of cross-talk due to PKP triplication
270	is also valid for the suppression of source-side reverberations – due to differences in

- 271 the source depths of the different earthquakes, the cross-talk in the autocorrelation
- 272 between the transmission and the source-side reverberation would be suppressed when

273 summing over the different earthquakes due to destructively interference (Draganov et

274 al., 2004, 2006).

275

276 Predictive deconvolution and seismic migration

277The bottom of the sedimentary basin (top of basement) often causes 278relatively strong free-surface multiples (Hansen and Johnson, 1948). The depth of the 279Malargüe basin (a sub-basin in the Neuquén basin) below the T-array is known 280(Nishitsuji et al., 2014). This allows us to suppress the basement free-surface multiples 281by applying a predictive-deconvolution filter (Yilmaz, 1987) based on the estimated 282two-way traveltime of these multiples. Note that such a filter was not used for the 283PV-array, as it is not above a basin (Moscoso et al., 2011). After interpreting the Moho 284below each subarray following as guidance the interpretation by Gilbert et al. (2006), 285we also apply predictive-deconvolution filter for possible free-surface multiples from 286the Moho. 287 As the subsurface structures might not be planar below the subarrays,

288 migration processing would be effective in moving dipping structures to their correct

289	location given an array has a sufficient length. In this study, we apply Kirchhoff
290	post-stack time migration (Yilmaz, 1987) to the GloPSI sections from the TN- and
291	TE-array. Migration is not applied for the PV-array due to its limited aperture; instead,
292	the individual traces are stacked.
293	As final processing steps, we apply lateral smoothing along the array to aid
294	the interpretation, using smoothed discretized splines based on the generalized
295	cross-validation (Garcia, 2010) (Figure 10), and then convert the migrated or stacked
296	traces from time to depth (Figure 11). For the depth conversion, we use a regional
297	velocity model down to 70 km depth (Gilbert et al., 2006) and the ak135 model
298	(Kennett et al., 1995) deeper than 70 km.
299	In Figure 10, we show a comparison of the obtained images when source
300	time functions of 10 s and 12 s are used in the estimation of $\overline{E}_n(t)$. It can be seen that
301	the different values give comparable results, which shows the robustness of the
302	procedure. The only substantial difference between the images in Figure 10 is in the
303	interpretation of the top of Moho. When using a two-way traveltime of 12 s, it seems
304	that the Moho is largely removed due to its consistent depth over the subarrays.

305	Although it might be possible to improve the time window by taking into account
306	individual source time functions, we found that the constant time window of 10 s is
307	sufficiently effective as we do not see major differences with the result when using a
308	window of 12 s. According to Kanamori and Brodsky (2004), the time window of 10 s
309	covers source time functions for earthquakes smaller or equal to $M_{\rm W}6.5$. Only 8% of
310	the earthquakes used for the TN array has $M_{\rm W} > 6.5$.
311	For the GO05 station, we apply the same processing as for the PV-array,
312	except that during the depth conversion we apply the velocity model as used for the
313	C02A station of the Talca Seismic Network in Dannowski et al. (2013) who utilized the
314	velocity model of Bohm et al. (2002). An approximation of $\overline{E}_n(t)$ is calculated by
315	taking the average of the retrieved results for GO05 and stations GO04 and GO06,
316	which are the N-S neighbors of GO05 in the Chilean National Seismic Network.
317	
318	Ouality control of the results at the seismic zone of the Nazca slab

319 For quality-control purpose, we first apply GloPSI to station GO05, which is 320 situated above the seismic zone of the slab. In the processed traces, the peak and

321	trough of the wiggles correspond to depths of P-wave impedance contrasts. We
322	compare the obtained GloPSI zero-offset reflection trace with the receiver-function
323	trace obtained for C02A in Dannowski et al. (2013), see Figure 11a. From the
324	receiver-function results, Dannowski et al. (2013) estimate the Moho depth at this
325	location at 33 km. GloPSI for GO05 also shows strong amplitude around 33 km
326	(Figure 11a). Note that around this depth starts a cluster of hypocenters (Figures 2 and
327	11a). Hypocenter clustering delineates the slab, meaning that beneath GO05 the strong
328	positive peaks at depths of about 40 km and 70 km correspond to the slab's top and
329	bottom, respectively (dashed green lines in Figure 11a). The correspondence of the
330	imaged reflectivity with the hypocenter clustering, but also with the slab's bottom from
331	the receiver-function trace (second positive peak at C02A trace in Figure 11a) confirms
332	the validity of applying GloPSI for slab imaging. Imaging reflectivity that is as strong
333	as the Moho means, that below GO05 the slab is locally (nearly) flat (Figures 1a and
334	1b). If the slab were locally inclined, the image would have exhibited lack of
335	reflectivity (Figure 1c).
336	

337 **Results Interpretation and Discussion**

338 Aseismic spot beneath the Peteroa volcano (PV-array)

339	Similar to the trace for station GO05, beneath the PV-array GloPSI reveals
340	the Moho where the strongest amplitude is seen, that is at a depth of about 45 km
341	(Figure 11b). This depth shows good agreement with a recent result of Gravity field
342	and Ocean Circulation Explorer (GOCE) operated by European Space Agency (ESA,
343	www.ea.int/ESA) (e.g., Reguzzoni et al., 2013) that shows the Moho depth to be
344	around 45 km in this region. A feature further down in the zero-offset reflection trace
345	from the PV-array is the appearance of reflectivity packages at around 100 km and 150
346	km depth, where the hypocenters of some intermediate-depth earthquake are present
347	(Figure 11b). Another striking feature is the lack of reflectivity for about 15 km around
348	the depth of 125 km. The latter corresponds to an aseismic spot at the Nazca slab.
349	Because of the aseismicity and because GloPSI would not image structures where no
350	impedance contrast exists (after applying predictive-deconvolution filter for possible
351	free-surface multiples from the Moho), the lack of reflectivity might be interpreted as
352	caused by certain amount of melt. If melted substance is indeed present around 125 km

353	depth, then one possible interpretation of the two strong-reflectivity packages at 100
354	km and 150 km depth would be as reflections from slab deformation, which in turn
355	would be caused by the melted substance. The deformation might be in the form of
356	detachment, shearing, necking, or any combination thereof. We illustrate the three pure
357	deformation scenarios in Figure 11d. The present hypocenters indicate vaguely the
358	slab, which is generally characterized as steeply dipping in this zone. The dip would be
359	too steep to retrieve reflections of a dipping interface delineating the slab (Figure 1c),
360	but deformations at the slab would give rise to scattered energy. Some of this energy
361	will be in the form of (nearly) vertically scattered fields, which will be recorded at the
362	station (Figure 1d). The latter will be turned by GloPSI into zero-offset reflections, and
363	consecutively imaged. If the slab is indeed deformed, depending on its thickness (e.g.,
364	the transparent green ellipses in Figure 11d), the primary reflection from the top of the
365	slab on one side of the deformation might interfere with the primary reflection from the
366	bottom of the slab from the other side of the deformation, which would make the
367	interpretation of the exact limits of the slab ambiguous. Because of this, in Figure 11b
368	we indicate with dashed green lines only the extent of the possible deformation of the

369 slab. We interpret the bottom of the slab at around 175 km.

370	Note that if melt is present and forms an impedance contrast with the mantle
371	and/or the slab, GloPSI would retrieve a reflection from this contrast as well unless the
372	melt itself forms a steeply dipping structure (Yilmaz, 1987). However, if there is no or
373	only weak impedance contrast due to, for example, the gabbro-eclogite transformation
374	of the slab, GloPSI will not retrieve a clear reflection from the melt. Frank et al. (2014)
375	showed that SI could be applied to S-wave phases as well (e.g., S, SS, ScS, and SKS).
376	S-waves have the advantage that they are more sensitive to melt than P-waves and thus
377	can provide extra information. An implementation of GloPSI to S-wave phases would
378	entail the use of global phases like PKS and SKS. Such implementation to our
379	temporary deployment would be challenging due to the low signal-to-noise ratio on the
380	horizontal components and the attenuation of much of the S-wave phases below the
381	sensitivity bandwidth of the instruments.
382	We do not exclude other possible interpretations for the lack of reflectivity
383	around 125 km. However, our interpretation is a logical consequence of the presence of
384	only a few intermediate-depth earthquakes: the slab here is insufficiently brittle to

385	generate many earthquakes and that might be indicative of a presence of magma with
386	possible slab deformation. Our interpretation is in a good agreement with results from
387	recent geochemical investigations of Jacques et al. (2013) suggesting that the
388	Planchón-Peteroa complex erupts not only lithospheric magma from the heterogeneous
389	mantle, but also magma from the Nazca slab.
390	

391 Aseismic zone of the Nazca slab beneath the T-array

392 The migrated images obtained from the results retrieved from GloPSI 393 beneath the TN- and TE-arrays are shown in Figure 11c. With the receiver-function 394 method, Gilbert et al. (2006) interpreted an apparently bifurcated Moho, with possibly 395a magma chamber in between, to be present in this region. Our result shows two strong 396 positive peaks, which appears to confirm the observation of Gilbert et al. (2006). Based 397 on their interpretation, we label the Moho and the magma chamber in Figure 11c where 398 the trough in blue is imaged at a depth of about 40 km. Our GloPSI image shows that 399 the bifurcation is continuous beneath the TN-array, but wedges out to the east beneath 400 the TE-array.

401	The image of the upper mantle beneath both arms of the T-array reveals a
402	complex structure. This heterogeneous image might correspond to the interpretation of
403	the study of Jacques et al. (2013). In their study, the authors indicated that the mantle
404	wedge in this region seems to be characterized, from a point of view of geochemical
405	components, by crustal assimilation or mantle heterogeneity. Note that if non-primary
406	reflections and spurious phases from autocorrelation cross-talk are retrieved, they will
407	contribute to the apparent complexity of the structure. The latter could be caused by
408	source-side reflections (even though we expect such cross-talk to be suppressed by the
409	summation over the different earthquakes), micro-seismic noise, etc.
409 410	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged
409 410 411	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged reflectors, indicated by the dashed green line in Figure 11c. This discontinuity is
409410411412	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged reflectors, indicated by the dashed green line in Figure 11c. This discontinuity is clearly observed below the TE-array from the middle of the array (100 km depth)
 409 410 411 412 413 	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged reflectors, indicated by the dashed green line in Figure 11c. This discontinuity is clearly observed below the TE-array from the middle of the array (100 km depth) towards the east (150 km depth). Due to the limited aperture of the T-array, deeper
 409 410 411 412 413 414 	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged reflectors, indicated by the dashed green line in Figure 11c. This discontinuity is clearly observed below the TE-array from the middle of the array (100 km depth) towards the east (150 km depth). Due to the limited aperture of the T-array, deeper steeply dipping structures will not be imaged, but will manifest themselves as lack of
 409 410 411 412 413 414 415 	summation over the different earthquakes), micro-seismic noise, etc. Below 100 km, we notice a pronounced discontinuity of the imaged reflectors, indicated by the dashed green line in Figure 11c. This discontinuity is clearly observed below the TE-array from the middle of the array (100 km depth) towards the east (150 km depth). Due to the limited aperture of the T-array, deeper steeply dipping structures will not be imaged, but will manifest themselves as lack of reflectivity (Figure 4-43 in Yilmaz, 1987). For instance, to record the free-surface

417	characterized by a dip of 40° and depth of 200 km, we need a receiver at the free
418	surface with an offset from the virtual-source position of more than 1000 km (Figure
419	1c). This can also be said in another way: to retrieve zero-offset reflection from a
420	structure with a dip of 40°, we will need to record incoming phases with incidence
421	angle of 40° as well, which is not possible with global phases. Although some
422	reflection discontinuities may be seen shallower than 150 km, it is difficult to interpret
423	them without other geophysical information. Note that a longer seismic array would be
424	required to better interpret the mantle structure. Since there is a possible remnant of an
425	upwelling plume in this region (Burd et al., 2014), some of these discontinuities might
426	be related to the plume, but they might also be related to a part of the mantle
427	convection or partial melting.
428	Let us look at the deeper part of the GloPSI image, where, based on the
429	extrapolation of the mapped hypocenters, we expect to see the Nazca slab. A
430	dimmed-reflectivity zone (between the dashed green lines) is visible beneath the
431	TN-array dipping from NNW around a depth of 180 km to 200 km to the SSE. This

432 zone causes discontinuity in the strong laterally coherent horizons A and B in Figure

433	11c. Beneath the TE-array, the GloPSI image exhibits a clear dimmed-reflectivity zone
434	(between the dashed green lines) dipping with an angle of 43° to the east and causing
435	discontinuity in horizon B. Note that horizon B is also visible around 62.5 s in Figure
436	10. The dimmed reflectivity might be caused by lack of impedance contrasts. This,
437	though, would not result in discontinuity of the imaged reflectors. As explained above,
438	another reason for the dimmed reflectivity might be the presence of dipping reflectors,
439	which, because of their depth and the relatively short array length, would not be well
440	imaged in the (migrated) section (Yilmaz, 1987). The presence of such dipping
441	reflectors would be manifested by discontinuity in horizontal reflectors (Figure 11c).
442	That is why, we interpret this dipping dimmed-reflectivity zone as the top and bottom
443	of the aseismic zone of the Nazca slab. We see that this part of the interpreted slab is
444	continuous and that the reflectivity does not indicate a possible slab deformation at this
445	latitude (35.5°S). Since there is no seismicity along this part of the slab, the condition
446	of this steeply dipping slab zone might be different from the condition in the shallower
447	zone where seismicity is present. This might support the interpretation of Yuan et al.
448	(2000) who proposed a completion of the eclogite transformation along this part of the

449 slab.

450

451 **Conclusions**

452We presented seismic interferometry with global phases (GloPSI) for 453imaging the aseismic and seismic parts of a subducting slab and the mantle above it. 454 GloPSI retrieves reflection responses from coinciding virtual source and receiver at 455each seismic station to which it is applied. We applied the method to global P-wave 456 phases recorded by an array of short-period stations installed for one year in the 457Malargüe region, Argentina, located east of the southern part of central Chile. The 458array consisted of a station distribution to the east of the Peteroa volcano and two linear subarrays to the east of the town of Malargüe. We processed the retrieved 459460 reflection responses to obtain depth images of the subsurface beneath the array. The images to the east of Malargüe town revealed, with high horizontal and vertical 461 462 resolution, a bifurcated Moho and a complex-structured upper mantle. On the images, 463 we also interpreted the aseismic part of the Nazca slab, which manifested itself as 464 dimmed reflectivity due to the relation between the depth of the dipping reflectors and

465	the short array length we used. The aseismic part of the slab appears to be without tears
466	and to be dipping with an angle of 43° to the east. The image beneath Peteroa also
467	showed the Moho. The deeper part of the image shows packages of strong reflectivity
468	with lack of reflectivity between them. These might be interpreted as a deformation in
469	the dipping slab. If so, the interpreted deformation could be in the form of detachment,
470	shearing, necking, or any combination thereof.
471	

472 Acknowledgements

473The data used in this study are collected using Java version of Windows Extracted 474from Event Data (JWEED) of Incorporated Research Institutions for Seismology 475(IRIS, http://www.iris.edu/dms/nodes/dmc/) and a reference earthquake catalogue from 476 U.S. Geological Survey (USGS, http://earthquake.usgs.gov/earthquakes/). This 477research is supported by the Division for Earth and Life Sciences (ALW) with financial 478aid from the Netherlands Organization for Scientific Research (NWO) with grant VIDI 479 864.11.009. The research of E.R. and K.W. was supported by the Netherlands Research 480 Centre for Integrated Solid Earth Sciences (ISES). The authors thank IRIS-PASSCAL

481	for providing the seismic equipment and the Argentine Ministry of Science,
482	Technology and Production Innovation for the financial support connected to the
483	transportation of the equipment. The authors also thank Pierre Auger Observatory and
484	the department of Civil Defense of Malargüe for the help during the data acquisition.
485	The authors are thankful to Issei Doi at Kyoto University for a discussion. The authors
486	thank Nori Nakata at Stanford University for his constructive comments on an earlier
487	version of the manuscript. The authors also thank Sjoerd de Ridder, the assistant
488	special editor, and two anonymous reviewers for their constructive comments that
489	improved the quality of this manuscript. The maps were drawn with
490	GenericMappingTool (GMT) (Wessel and Smith, 1991).
491	

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645

646 Figure captions

647	Figure 1. : A schematic illustration of how GloPSI would or would not retrieve
648	reflection responses for: (a) a horizontally layered structure and vertical
649	transmission responses; (b) a gently dipping structure and nearly vertical
650	transmission responses; (c) as in (b), but for a steeply dipping structure; (d)
651	as in (c), but when an abrupt change (e.g., slab deformation) presents in
652	present in the lateral continuation of the dipping structure. The black lines
653	indicate the transmission response from the global earthquakes, while the
654	gray dashed lines depict the reflection response that will not be recorded at

655	the	station	due	to	the	configuration.	Two-way	arrows	indicate	the
656	refl	ection re	spons	se th	at w	ill be recorded a	t the station	1.		

- Figure 2. : Center Location of the seismic stations used in our study, and hypocenters
- mapping using earthquakes archived by USGS. Below and right distribution
 of the hypocenters in depth within the red dashed-line areas in NWW-SEE and
 NNE-SSW direction.
- Figure 3. : Distribution of the global-phase earthquakes used in our study. The circlesshow the location of the earthquakes used for MalARRgue and the GO05
- station. The location of MalARRgue is indicated by the black triangle with
- its topography maps (Becker et al., 2009) in the insets. The distribution of
- the back azimuth of the earthquakes for the T-array is shown in the inset.
- 666 Figure 4. : An example recording of a global earthquake on the vertical component of
- 667 the stations from the TN-array. The area highlighted in light blue indicates668 the used window that contains the global phases. The orange and green
- 669 lines indicate the P- and S-wave phase onsets by IRIS, respectively.

670 Figure 5. : The computed power spectral densities for four earthquakes with different

671	magnitudes that occurred at global distances. The densities are computed
672	for station TE01 of the TE-array in MalARRgue. Δ indicates the
673	epicentral distances of the global earthquakes.
674	Figure 6. : Number of original and interpolated global phases for TN- (top) and
675	TE-array (bottom) stations.
676	Figure 7. : GloPSI results retrieved at the MalARRgue stations before seismic
677	processing. The annotations along the horizontal axis show the actual
678	station codes.
679	Figure 8. : The results from Figure 7 after subtraction of the mean $\overline{E}_n(t)$ per subarray.
680	Figure 9. : A comparison of GloPSI images obtained when including and when
681	excluding global phases with PKP triplications. The number of
682	earthquakes for the TN(TE)-array with and without the PKP triplications
683	are 72 (85) and 13 (13), respectively.
684	Figure 10. : GloPSI results for the TN- and TE-array after post-stack time migration
685	with lateral smoothing in the offset orientation when respective source
686	time functions of 10 s and of 12 s are used in the estimation of $\overline{E}_n(t)$.

687	Figure 11. : Summarized interpretation with seismicity along the NWW-SEE area of
688	GloPSI for MalARRgue and station GO05. a. GloPSI for GO05 and
689	receiver function for C02A at the Nazca-slab seismic zone. Moho depth is
690	interpreted using receiver function (modified from Dannowski et al., 2013)
691	at C02A. b. GloPSI for the PV-array beneath the Peteroa Volcano. c.
692	GloPSI for the TN- and TE-array at the Nazca-slab aseismic zone. Dashed
693	green lines in the panels indicate where we interpret the Nazca slab and
694	transparent green rectangles indicate possible interval of the interpretation.
695	The transparent green ellipses indicate where we interpret the Nazca-slab
696	deformation, while the transparent gray triangle – the possible connection
697	between the Nazca-slab seismic and aseismic zones in three dimensions.
698	The insets in the bottom left corner illustrate three possible scenarios
699	explaining the retrieved strong reflectivity below the PV-array. Gray
700	circles (some transparent for visibility purposes) indicate earthquake
701	hypocenters.

Table 1. Glob	al-phase seis	smic used	in this st	udy	14 A
Date (month/d/ur)	Time (hr:min:c)	Lat.	Lon.	Dep.	M _W Array ID
01/18/12	12:50:21	-0.877	126.829	(KIII) 19	5 7 TF
01/28/12	0:17:11	13.386	124.586	35	5.5 TE
02/04/12	13:09:23	11.872	125.754	12	5.8 PV/TN/TE/GO
02/06/12	3:49:13	9.999	123.206	11	6.7 TE
02/06/12	4:20:00	10.092	123.227	10	5.6 TE
02/06/12	10:10:20	9.885	123.095	15	5.0 PV/1N/1E/GO
02/06/12	6:22:01	36 214	141 386	28	5.8 PV/TN/TE/GO
02/26/12	2:35:01	22.661	120.891	28	5.9 TE
02/26/12	6:17:20	51.708	95.991	12	6.6 PV/TN/TE/GO
02/29/12	14:32:48	35.200	141.001	26	5.6 TE
03/08/12	22:50:08	39.383	81.307	38	5.9 TE
03/12/12	6:06:41	36.741	73.152	11	5.7 PV/IN/IE
03/12/12	9:08:35	45.239	147.009	110	5.0 PV/IN/IE 6.9 PV/TN/TE
03/14/12	10:49:25	40.881	144 761	10	6.1 PV/TN/TE
03/14/12	12:05:05	35.687	140.695	10	6.0 PV/TN/TE
03/16/12	7:58:02	10.037	125.633	18	5.8 PV/TN/TE/GO
03/22/12	0:21:37	3.513	125.859	117	5.6 TE
03/27/12	11:00:45	39.859	142.017	15	6.0 PV/TN/TE/GO
04/01/12	14:04:25	37.110	140.957	48	5.8 PV/1N/1E
04/11/12	0.30.57	0.802	92 463	20	8.0 FV/TN/TE/GO 8.2 PV/TN/TE/GO
04/13/12	10:10:01	36 988	141 152	11	5.7 PV/TN/TE/GO
04/14/12	15:13:14	49.380	155.651	90	5.6 TE
04/15/12	5:57:40	2.581	90.269	25	6.3 PV/TN/TE/GO
04/20/12	22:19:47	3.256	93.853	25	5.8 TE
04/20/12	22:28:59	3.269	93.821	22	5.8 PV/TN/TE/GO
04/20/12	23:14:31	2.158	93.360	28	5.9 PV/TN/TE/GO
04/21/12	1:10:55	-1.01/	134.276	10	0.7 PV/1N/1E/GO 5.7 TE/GO
04/23/12	21:21:45	48.397	154.739	46	5.7 PV/TN/TE
04/24/12	14:57:10	8.868	93.949	14	5.6 PV/TN/TE/GO
04/25/12	7:42:23	9.011	93.945	9	5.9 PV/TN/TE/GO
04/29/12	8:09:04	2.704	94.509	14	5.7 PV/TN/TE/GO
04/29/12	10:28:52	35.596	140.349	44	5.8 PV/TN/TE/GO
05/12/12	23:28:44	38.612	70.354	10	5.7 PV/IN/IE/GO
05/25/12	10.02.25	41.555	142.082	40	6.1 PV/TN/TE
06/09/12	14:23:20	48.851	154.852	49	5.5 TE
06/09/12	21:00:18	24.572	122.248	70	5.9 PV/TN/TE
06/11/12	5:29:12	36.023	69.351	16	5.7 TE
06/14/12	20:17:25	1.293	126.828	61	5.5 TE
06/15/12	1:14:08	5.719	126.354	41	5.7 PV/TN/TE/GO
06/16/12	22:18:47	15.593	119.563	28	5.9 PV/1N/1E/GO
06/23/12	4.34.53	3 009	97 896	95	6.1 PV/TN/TE/GO
06/29/12	21:07:34	43.433	84.700	18	6.3 PV/TN/TE/GO
07/08/12	11:33:03	45.497	151.288	20	6.0 PV/TN/TE
07/11/12	2:31:17	45.401	151.424	10	5.7 PV/TN/TE
07/12/12	12:51:59	45.452	151.665	12	5.7 TE
07/12/12	14:00:34	36.527	70.906	198	5.8 PV/TN/TE
07/19/12	2:40:12	37.248	/1.3/3	98	5.6 PV/IN/IE/GU
07/20/12	6:10:25	49.407	155.907	19	6.0 PV/TN/TE/GO
07/20/12	6:32:56	49.354	156.132	10	5.9 PV/TN/GO
07/25/12	0:27:45	2.707	96.045	22	6.4 PV/TN/GO
08/11/12	12:23:18	38.329	46.826	11	6.5 TE
08/11/12	12:34:36	38.389	46.745	12	6.4 TE
08/12/12	2:50:28	35.661	82.518	592	6.2 PV/IN/IE/GO 7.7 PV/TN/TE
08/18/12	9:41:52	-1.315	120 096	10	6.3 PV/TN/TE
08/18/12	15:31:40	2.645	128.697	10	5.8 TE
08/25/12	14:16:17	42.419	142.913	55	5.9 PV/TN/TE/GO
08/26/12	15:05:37	2.190	126.837	91	6.6 PV/TN/TE/GO
08/29/12	19:05:11	38.425	141.814	47	5.5 PV/TN/TE/GO
08/31/12	12:47:55	10.811	120.038	28	5.6 PV/TN/TE/CO
09/03/12	6:49:50	6.610	123.875	12	5.9 PV/TN/TE/GO
09/03/12	18:23:05	-10.708	113.931	14	6.3 PV/TN/GO
09/03/12	19:44:22	7.905	125.044	10	5.7 PV/TN/TE/GO
09/08/12	6:54:19	21.527	145.923	5	5.6 TE
09/08/12	10:51:44	-3.177	135.109	21	6.1 PV/TN/GO
09/09/12	5:39:37	49.247	155.750	31	5.9 IE 5.5 PV/TN/TE/CO
09/11/12	16:36:50	11 838	143 218	14 8	5.9 TE
09/14/12	4:51:47	-3.319	100.594	19	6.3 PV/TN/GO
10/01/12	22:21:46	39.808	143.099	15	6.0 PV/TN
10/08/12	11:43:31	-4.472	129.129	10	6.2 PV/TN/GO
10/12/12	0:31:28	-4.892	134.030	13	6.6 PV/TN/GO
10/14/12	9:41:59	48.308	154.428	35	5.8 PV/TN 5.6 PV/TN
10/10/12	4-42-30	49.018	120.438	326	5.0 F V/1N 6.0 PV/TN
11/01/12	23:37:18	1.229	122.105	35	5.5 TE
11/02/12	18:17:33	9.219	126.161	37	6.1 TN/TE/GO
11/05/12	4:30:27	37.791	143.610	19	5.6 TN/TE/GO
11/06/12	1:36:22	1.374	122.200	25	5.6 TN/TE/GO
11/06/12	1:42:26	1.357	122.167	35	5.6 TE
11/11/12	1:12:39	23.005	95.885	14	6.8 TN/TE/GO
11/14/12	3.21:42	9.982 40.280	122.472	41 20	6.5 TN/TE/GO
11/27/12	7:34:25	17.684	145.763	192	5.5 TE
12/07/12	8:18:23	37.890	143.949	31	7.3 PV/TN/TE/GO
12/09/12	21:45:35	6.703	126.166	63	5.8 PV/TN/TE/GO
12/10/12	16:53:09	-6.533	129.825	155	7.1 PV/TN/GO
12/11/12	6:18:27	0.533	126.231	30	6.0 PV/TN/TE

Date, Time, Lat., Lon., Dep. and M_w , the moment magnitude, are provided by USGS (http://earthquake.usgs.gov/earthquakes/). For Array

ID, PV, TE, TN, and GO indicate PV-array, TE-array, TN-array, and GO05, respectively.



Figure 1. : A schematic illustration of how GloPSI would or would not retrieve reflection responses for: (a) a horizontally layered structure and vertical transmission responses; (b) a gently dipping structure and nearly vertical transmission responses; (c) as in (b), but for a steeply dipping structure; (d) as in (c), but when an abrupt change (e.g., slab deformation) presents in present in the lateral continuation of the dipping structure. The black lines indicate the transmission response from the global earthquakes, while the gray dashed lines depict the reflection response that will not be recorded at the station due to the configuration. Two-way arrows indicate the reflection response that will be recorded at the station. 124x115mm (300 x 300 DPI)



Figure 2. : Center – Location of the seismic stations used in our study, and hypocenters mapping using earthquakes archived by USGS. Below and right – distribution of the hypocenters in depth within the red dashed-line areas in NWW-SEE and NNE-SSW direction. 153x116mm (300 x 300 DPI)



Figure 3. : Distribution of the global-phase earthquakes used in our study. The circles show the location of the earthquakes used for MalARRgue and the GO05 station. The location of MalARRgue is indicated by the black triangle with its topography maps (Becker et al., 2009) in the insets. The distribution of the back azimuth of the earthquakes for the T-array is shown in the inset. 149x82mm (300 x 300 DPI)



An example recording of a global earthquake on the vertical component of the stations from the TN-array. The area highlighted in light blue indicates the used window that contains the global phases. The orange and green lines indicate the P- and S-wave phase onsets by IRIS, respectively. 233x125mm (300 x 300 DPI)



The computed power spectral densities for four earthquakes with different magnitudes that occurred at global distances. The densities are computed for station TE01 of the TE-array in MalARRgue. Δ indicates the epicentral distances of the global earthquakes. 173x246mm (300 x 300 DPI)



Figure 6. : Number of original and interpolated global phases for TN- (top) and TE-array (bottom) stations. 279x361mm (300 x 300 DPI)



Figure 7. : GloPSI results retrieved at the MalARRgue stations before seismic processing. The annotations along the horizontal axis show the actual station codes. 215x166mm (300 x 300 DPI)



Figure 8. : The results from Figure 7 after subtraction of the mean averaged source time function per subarray. 215x166mm (300 x 300 DPI)



Figure 9. : A comparison of GloPSI images obtained when including and when excluding global phases with PKP triplications. The number of earthquakes for the TN(TE)-array with and without the PKP triplications are 72 (85) and 13 (13), respectively. 215x166mm (300 x 300 DPI)

Figure 10. : GloPSI results for the TN- and TE-array after post-stack time migration with lateral smoothing in the offset orientation when respective source time functions of 10 s and of 12 s are used in the estimation of the averaged source time function. 179x120mm (300 x 300 DPI)

Figure 11. : Summarized interpretation with seismicity along the NWW-SEE area of GloPSI for MalARRgue and station GO05. a. GloPSI for GO05 and receiver function for C02A at the Nazca-slab seismic zone. Moho depth is interpreted using receiver function (Dannowski et al., 2013) at C02A. b. GloPSI for the PV-array beneath the Peteroa Volcano. c. GloPSI for the TN- and TE-array at the Nazca-slab aseismic zone. Dashed green lines in the panels indicate where we interpret the Nazca slab and transparent green rectangles indicate possible interval of the interpretation. The transparent green ellipses indicate where we interpret the Nazca-slab deformation, while the transparent gray triangle – the possible connection between the Nazca-slab seismic and aseismic zones in three dimensions. The insets in the bottom left corner illustrate three possible scenarios explaining the retrieved strong reflectivity below the PV-array. Gray circles (some transparent for visibility purposes) indicate earthquake hypocenters.

206x176mm (300 x 300 DPI)